

Composition of cometary particles versus distance to sun during sample collection

- based on multivariate evaluation of mass spectral data (Rosetta/COSIMA)



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Rosetta was a space mission (ESA) working more than two years (Aug 2014 to Sep 2016) near comet Churyumov-Gerasimenko (short 67P).



COSIMA was an instrument on board of Rosetta. It collected cometary particles with size 10 - 1000 μm on metal targets (distance from the comet 10 - 1500 km). More than 30,000 particles were documented by images, and more than 35,000 secondary ion mass spectra (TOF-**SIMS**) were measured on about 250 particles [1 - 7].

Particle (sample) collection occurred at heliocentric distances (from comet to sun) between 1.8 AU and 3.8 AU (AU for Astronomical Unit, 150 million km, approximately the mean distance between earth and sun).

Aim of this work was to check relationships between the heliocentric distance of sampling and mass spectral data.

Data

- n = 2863 objects derived from TOF-SIMS spectra measured at 256 particles, collected in 412 sampling intervals $n_1 = 1459$ in class A: collection at ≤ 1.7 AU (*near* sun, warmer)
- n₂ = 1404 in class B: collection at >1.7 AU (far from sun, cooler)
- m = 9 variables: ion counts for C⁺, CH⁺, CH₂⁺, CH₃⁺, Mg⁺, C₂H₃⁺, $C_3H_3^+$, $C_3H_4^+$, Fe⁺, normalized to sum 100 or transformed by centered log-ratios (clr) [compositional data].



Strategy for Data Evaluation

O Characterization of the discrimination of class A (particles collected closer to sun) and class B (more far from sun).

O A good discrimination indicates a relationship between the used mass spectral data (characterizing the chemical composition) and the heliocentric distance of sampling.

O Methods applied: PCA; univariate distributions; KNN classification; PLS calibration; the last two with rdCV, repeated double cross validation [8].

Results



Conclusions

O The mass spectral data (characterizing the composition of the particles) are multivariate different for class A (collection near aphelion) and class B (collection more far from sun).

- O KNN separates A and B well; linear approaches (PCA, PLS) do not.
- **Q** Some ion count ratios show different distributions for A and B; single ions do not.
- O Influence of the distance to the comet or other parameters cannot be excluded.

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The instrument COSIMA [1] onboard of the ESA spacecraft Rosetta collected dust particles in the neighborhood of comet Churyumov-Gerasimenko [2]. The distance to the sun during the sample collections varied between 1.2 and 3.8 AU (AU for astronomical unit, 150 000 000 km). The chemical composition of the particle surfaces was characterized by COSIMA using TOF-SIMS (time-of-flight secondary ion mass spectrometry). A set of about 3000 spectra has been selected, and relative abundances for CH-containing positive ions (from organic compounds [3-6]) as well as elemental ions (from minerals [7]) define a set of multivariate data. Evaluation by chemometric techniques indicates different compositions of samples collected at different distances to the sun. The applied methods comprise (1) considering the compositional nature of the used mass spectral data and centered log-ratio transformation [8]; (2) robust PCA [9]; (3) KNN-classification with repeated double cross validation [10].

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- [1] Kissel J., et al.: Space Sci. Rev., 128, 823 (2007)
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